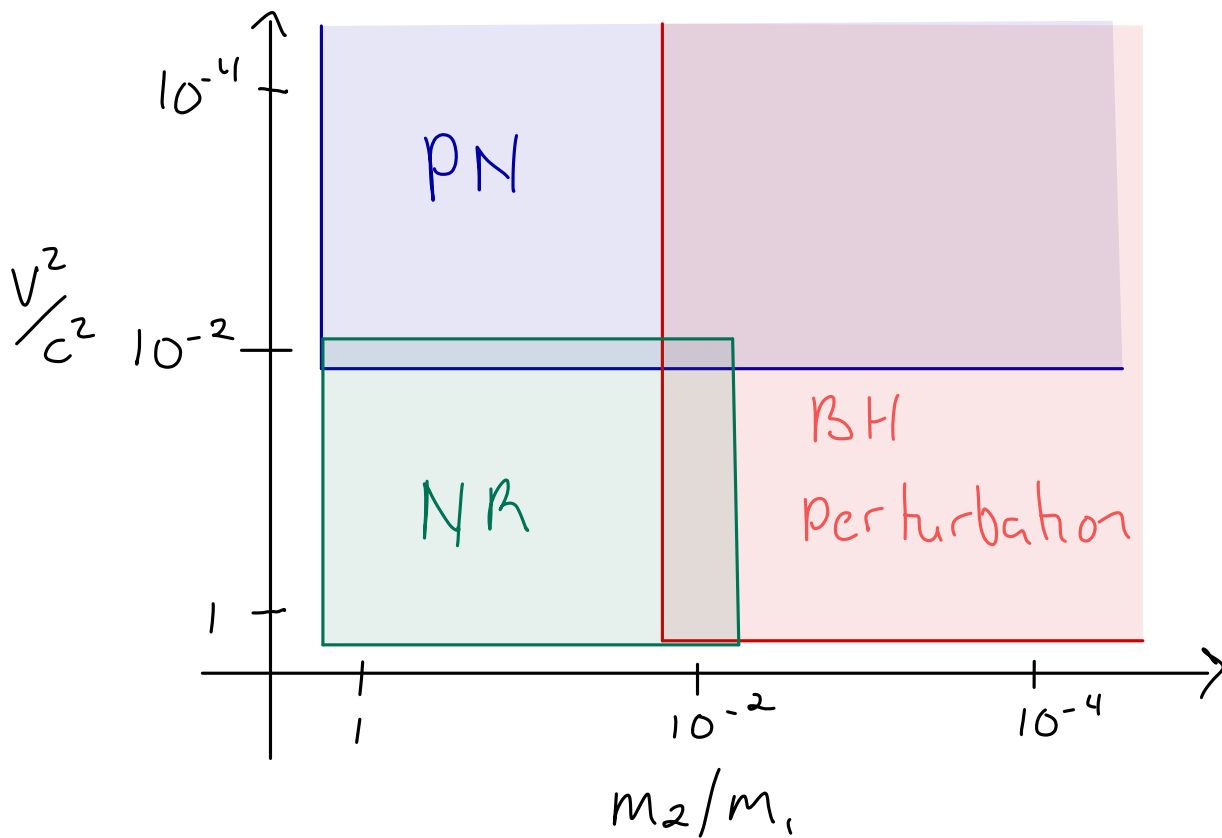


# Stationary Phase Approximation (SPA)

post-Newtonian theory: slow:  $(\frac{v}{c})^2 \ll 1$  Weak fields:  $\frac{GM}{rc^2} \ll 1$

Circular binary  $\Rightarrow$  relate  $r, v$  via  $\omega$

$$\left. \begin{aligned} \omega^2 &= \frac{GM}{r^3} + \dots \\ V &= \omega r \end{aligned} \right\} X \equiv (GM\omega)^{2/3} c^{-2} \left. \begin{aligned} \omega &= \frac{c^3}{GM} X^{3/2} \\ r &= \frac{GM}{c^2 X} + \dots \\ \frac{v^2}{c^2} &= X + \dots \end{aligned} \right\}$$

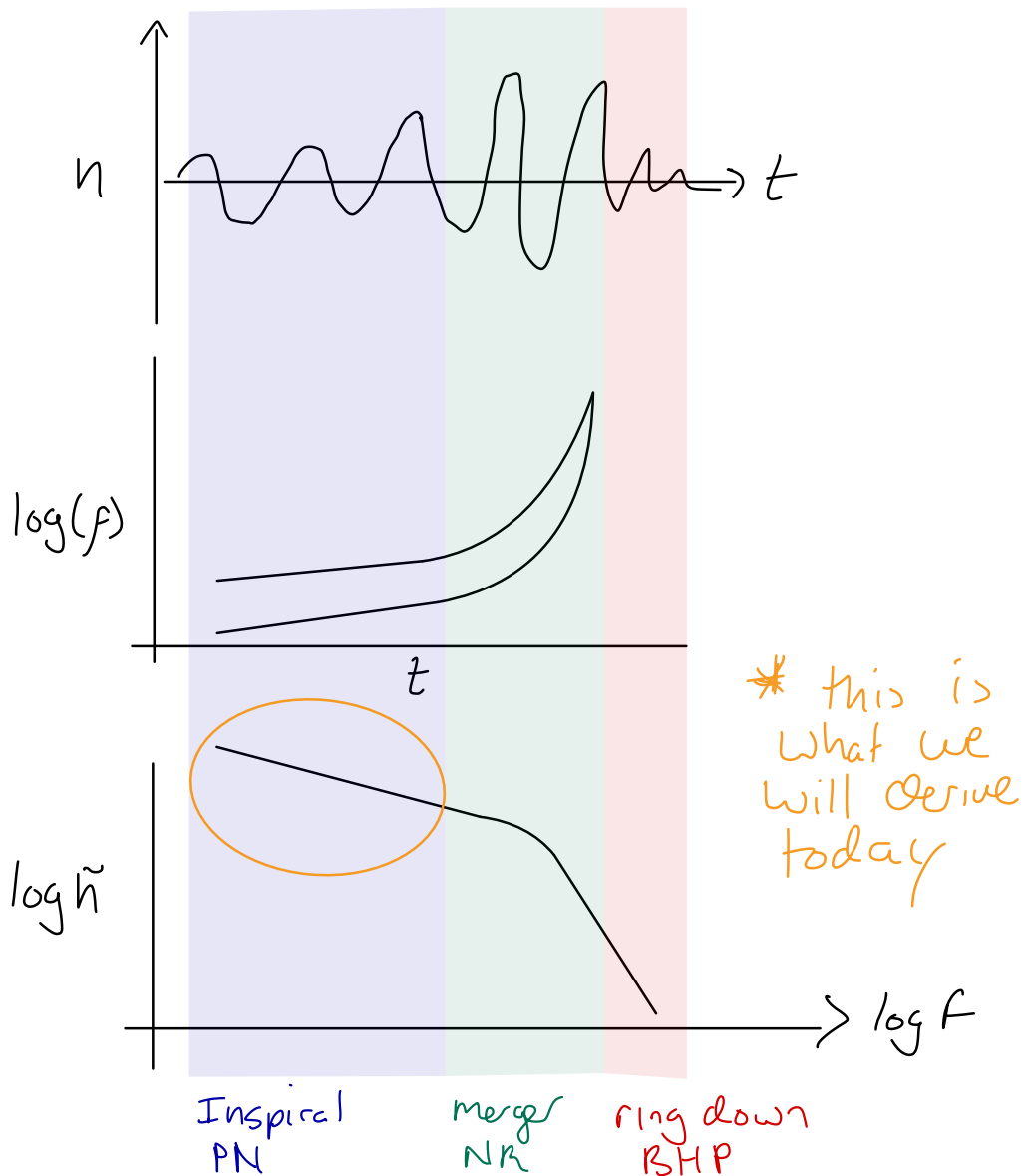


set  $G=c=1$  from now on...

We need  $\tilde{h}(f)$

1. signal identification (matched filtering)
2. parameter estimation

What does a gravitational wave signal look like?

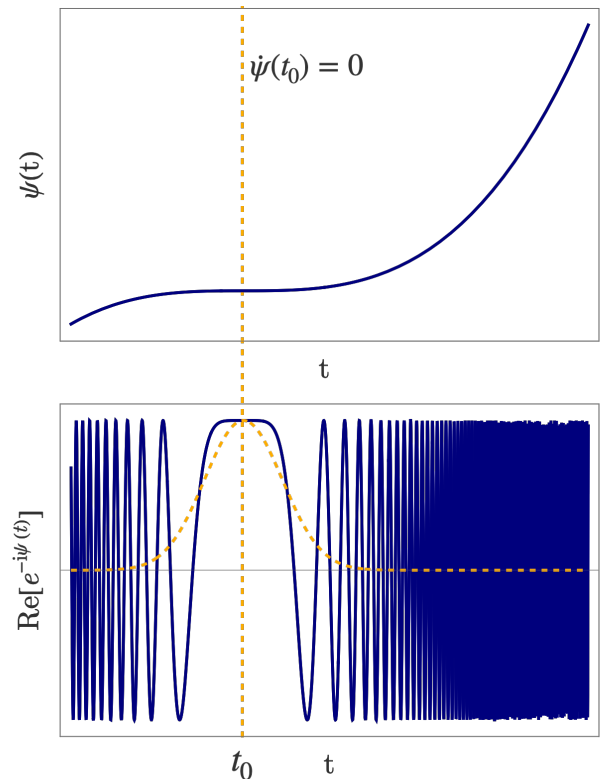


SPA  $\Rightarrow$  directly compute  $\vec{h}(f)$   
 when radiation-reaction time scale  $\ll$  orbital period

What do we need to apply it?

1. Amplitude varies slower than phase
2. orbital frequency / phase are positive and monotonically increasing

Resources: 0906.0313, SPA wikipedia, Maggiore, 0907.0700



First use SPA to compute the integral of a generic oscillating function

$$F(t) = A(t) e^{-i\psi(t)} \quad I = \int F(t) dt$$

If  $F(t)$  has a "stationary point"  $t_0$  where  $\psi(t_0) = 0$  the area near this point will contribute most to the integral

Expand  $A(t)$ ,  $\psi(t)$  about  $t_0$

$$\psi(t) = \psi(t_0) + \dot{\psi}(t_0)(t-t_0) + \frac{1}{2}\ddot{\psi}(t_0)(t-t_0)^2 + \dots$$

$$A(t) = A(t_0) + \dots \quad (\text{Amplitude varies slowly})$$

$$I = A(t_0) e^{-i\psi(t_0)} \int e^{-\frac{i}{2}\ddot{\psi}(t_0)(t-t_0)^2} dt \quad \text{Gaussian!}$$

$$I = A(t_0) \left[ \frac{2\pi}{\ddot{\psi}(t_0)} \right]^{1/2} e^{-i[\psi(t_0) + \pi/4]}$$

Now lets apply the SPA to obtain the frequency domain GW Strain

$$h^{ij} = h_+(t) \begin{pmatrix} 1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & 0 \end{pmatrix} + h_x(t) \begin{pmatrix} 0 & 1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

$$L=0$$

$$h_+ = \frac{1}{2}(1 + \cos^2 \iota) A(t) \cos[2\phi(t)] \quad h_x = \cos \iota A(t) \sin[2\phi(t)]$$

$\phi(t)$  = orbital phase

$$M = m_1 + m_2$$

$$h^{ij} = \frac{1}{d} \ddot{I}^{ij} + \dots \sim \frac{1}{d} \omega^2 m r^2$$

$$r = |\vec{r}_1 - \vec{r}_2|$$

$$\eta = \frac{m_1 m_2}{m^2} \quad m\eta = \mu$$

$$A(t) = -4 \frac{1}{d} \eta m \omega^2 r^2 + \dots$$

$$v = |\vec{v}_1 - \vec{v}_2|$$

$$A(t) = \boxed{-4 \frac{1}{d} \eta m \omega^2 r^2 + \dots}$$

While we are at it,

$$E = \frac{1}{2} m \eta v^2 - \frac{m^2 \eta}{r} = \boxed{-\frac{1}{2} m \eta \dot{x}^2 + \dots}$$

total energy

$$P \sim \ddot{I}^{ij} \ddot{I}_{ij} + \dots \sim \omega^6 (m r^2)^2$$

$$= \frac{32}{5} \eta^2 m^2 \omega^6 r^4 = \boxed{\frac{32}{5} \eta^2 \dot{x}^5 + \dots}$$

Power radiated by GWs

$$h(t) = h_r - i h_x = A e^{-2i\phi}$$

collect  $h_r/h_x$  into a single term for simplicity

$$\begin{aligned}\tilde{h}(f) &= \int h(t) e^{2\pi i f t} dt \\ &= \int A(t) e^{2\pi i f t} e^{-2i\phi(t)} dt\end{aligned}$$

We can write  $\tilde{h}(f) = \tilde{A}(f) e^{-i\tilde{\phi}(f)}$  in terms of an Amplitude and phase:

$$\tilde{A}(f) = A(t_0) \left[ \frac{2\pi}{\dot{\psi}(t_0)} \right]^{1/2} \quad \tilde{\phi}(f) = \psi(t_0) + \pi/4$$

the total phase of the integrand is

$$\psi(t) = -2 [\pi f t - \phi(t)] \quad \leftarrow \text{this carries a minus sign because of the way we defined } \psi(t) \text{ on the first page}$$


$$\dot{\psi}(t) = -2 [\pi f - \dot{\phi}(t)]$$

$$\ddot{\psi}(t) = 2 \dot{\phi}(t) = 2 \dot{\omega}(t)$$

$$\text{stationary point: } \dot{\psi}(t_0) = 0 \Rightarrow \dot{\phi}(t_0) = \pi f = \omega(t_0)$$

the Amplitude and phase of  $\tilde{h}(f)$  are

$$\tilde{A}(f) = A(t_0) \left( \frac{\pi}{\dot{\omega}(t_0)} \right)^{1/2}$$

 = we need to solve for these

$$\tilde{\phi}(f) = 2 t_0 \omega(t_0) - 2 \phi(t_0) - \pi/4$$

$$x_0 = x(t_0) = [m \omega(t_0)]^{2/3}$$

$$\tilde{\phi}(f) = -2 t_0 \frac{x_0^{3/2}}{m} + 2 \phi(t_0) + \pi/4$$

$$= (\pi m f)^{2/3}$$

Solve for  $\tilde{\Phi}(p)$  First

Balance equation  $P = -\frac{dE}{dt} = -\frac{dE}{dx} \frac{dx}{dt} = -E' \frac{dx}{dt}$

Not Newtonian! binary loses energy to GW

Find  $t_0$ :

$$dt = -\frac{E'}{P} dx$$

$$\frac{E'}{P} = -\frac{5}{64} \frac{m}{\eta} X^{-5}$$

$$t_0 = t_c - \int \frac{E'}{P} dx \Big|_{x=x_0}$$

$$t_c + \frac{5}{64} \frac{m}{\eta} \int X^{-5} dx \Big|_{x=x_0}$$

$$t_0 = t_c - \frac{5}{256} \frac{m}{\eta} X_0^{-4}$$

Find  $\phi(t_0)$

$$\omega = \frac{d\phi}{dt} = \frac{d\phi}{dx} \frac{dx}{dt} = -\frac{d\phi}{dx} \frac{P'}{E} = \frac{X^{3/2}}{m}$$

$$d\phi = -\frac{E'}{P} \frac{X^{3/2}}{m} dx$$

$$\phi(t_0) = \phi_c - \int \frac{E'}{P} \frac{X^{3/2}}{m} dx \Big|_{x=x_0}$$

$$= \phi_c + \frac{5}{64} \frac{1}{\eta} \int X^{-7/2} dx \Big|_{x=x_0}$$

$$\phi(t_0) = \phi_c - \frac{1}{32} \frac{1}{\eta} X_0^{-5/2}$$

$$\begin{aligned}
\tilde{\phi} &= -2t_c \frac{x_0^{3/2}}{m} + 2\phi(t_c) - \pi/4 \\
&= -2\left(t_c - \frac{5}{256} \frac{m}{\eta} x_0^{-4}\right) \frac{x_0^{3/2}}{m} + 2\left(\phi_c - \frac{1}{32} \frac{1}{\eta} x_0^{-5/2}\right) - \pi/4 \\
&= -2t_c \frac{x_0^{3/2}}{m} + 2\phi_c - \frac{3}{128} \frac{1}{\eta} x_0^{-5/2} + \pi/4
\end{aligned}$$

$$\tilde{\phi}(f) = -2\pi f t_c + 2\phi_c + \pi/4 - \frac{3}{128} \frac{1}{\eta} (\pi m f)^{-5/3}$$

$$\mathcal{M} = m \eta^{3/5}$$

$$\tilde{\phi}(f) = -2\pi f t_c + 2\phi_c + \pi/4 - \frac{3}{128} \mathcal{M}^{-5/3} (\pi f)^{-5/3}$$

chirp mass!

Next, solve for  $\tilde{A}(f)$

Find  $\dot{\omega}(t_0)$ :

$$\dot{\omega} = \frac{d\omega}{dx} \frac{dx}{dt}$$

$$\dot{\omega} = \frac{3}{2} \frac{1}{m} x^{1/2} \frac{dx}{dt}$$

$$\dot{\omega}(t_0) = \frac{3}{2} \frac{1}{m} x^{1/2} \frac{P}{E} \Big|_{x=x_0}$$

$$\dot{\omega}(t_0) = \frac{96}{5} \frac{\eta}{m^2} x_0^{11/2}$$

Find  $A(t_0)$ :

$$A(t_0) = -4 \frac{1}{2} \eta m x_0$$

$$\tilde{A}(f) = A(t_0) \left( \frac{\pi}{\dot{\omega}(t_0)} \right)^{1/2}$$

$$= -\left(\frac{5\pi}{6}\right)^{1/2} \frac{1}{2} m^2 \eta^{1/2} x^{-7/4}$$

$$\tilde{A}(f) = -\left(\frac{5\pi}{6}\right)^{1/2} \frac{1}{2} m^2 \eta^{1/3} (\pi m f)^{-7/6}$$

$$\tilde{A}(f) = -\left(\frac{5\pi}{6}\right)^{1/2} \frac{1}{2} \mathcal{M}^{5/6} (\pi f)^{-7/6}$$

Note:  $\tilde{h}_+ = \frac{1}{2} \tilde{h}$  ,  $\tilde{h}_x = \frac{i}{2} \tilde{h}$   $\Rightarrow \tilde{h} = \tilde{h}_+ - i \tilde{h}_x$

How to extend calculation?

must add corrections to:

Keplers law  $[w(r)]$ ,  $A(t)$ ,  $E(x)$ ,  $P(x)$

What we have computed is the Newtonian term of the Taylor-F2 waveform

see 2012.0135 for an overview of different waveform approximants